

Vectors (again)

function manifold
 $f: M \rightarrow \mathbb{R}$ is C^∞

a tangent space at $p \in M$ in the manifold
defines the directional derivatives

$f \rightarrow \frac{df}{dh}$ at p parameterized by h

if we parameterize by h & y
coordinates

$$\rightarrow x^h(h) \text{ & } x^y(y) \rightarrow a \frac{df}{dh} + b \frac{df}{dy}$$

the chain rule applies

$$\begin{aligned} \left(a \frac{d}{dh} + b \frac{d}{dy} \right) (fg) &= a \left(f \frac{dg}{dh} + g \frac{df}{dh} \right) + b \left(f \frac{dg}{dy} + g \frac{df}{dy} \right) \\ &= \left(a \frac{df}{dh} + b \frac{df}{dy} \right) g + \left(a \frac{dg}{dh} + b \frac{dg}{dy} \right) f \end{aligned}$$

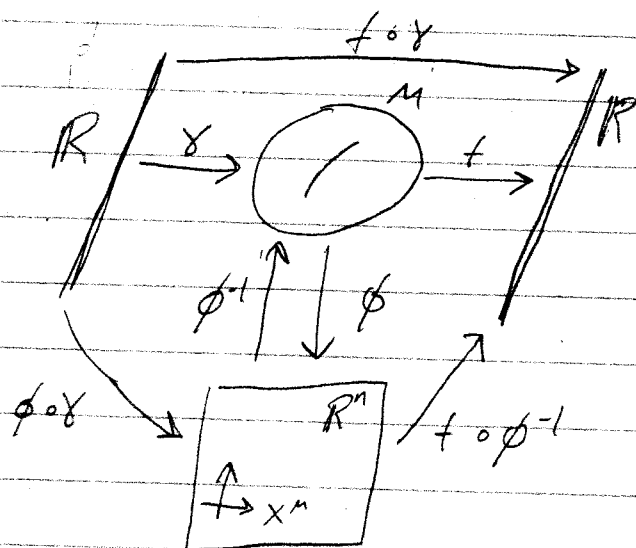
x^h - coordinate

$\{\partial_p\}$ - partial derivatives which form the basis of T_p

$\phi: M \rightarrow \mathbb{R}^n$ - coordinate chart

$\gamma: \mathbb{R} \rightarrow M$ - curve - parameterized by h

$f: M \rightarrow \mathbb{R}$ - function



$$\begin{aligned}
 \frac{d}{d\lambda} t &= \frac{d}{d\lambda} (t \circ \gamma) = \frac{d}{d\lambda} [(t \circ \phi^{-1}) \circ (\phi \circ \gamma)] \\
 &= \frac{d(\phi \circ \gamma)^m}{d\lambda} \frac{\partial (t \circ \phi^{-1})}{\partial x^m} \leftarrow \text{chain rule} \\
 &= \frac{dx^m}{d\lambda} \partial_m t
 \end{aligned}$$

$$\therefore \frac{d}{d\lambda} = \frac{dx^m}{d\lambda} \partial_m \quad \partial_m = \hat{e}_m$$

$\nwarrow$ coordinate basis for T_p

Note: coord basis vectors are not always normalized to unity or orthogonal, on a curved manifold a coord basis is never orthonormal throughout the neighborhood of a pt where curvature doesn't vanish

transformation law

$$\partial_{\mu'} = \frac{\partial x^\mu}{\partial x^{\mu'}} \partial_\mu$$

$$V^\mu \partial_\mu = V^{\mu'} \partial_{\mu'} = V^{\mu'} \frac{\partial x^\mu}{\partial x^{\mu'}} \partial_\mu$$

$$\therefore V^\mu = V^{\mu'} \frac{\partial x^\mu}{\partial x^{\mu'}}$$

$$V^{\mu'} = \frac{\partial x^{\mu'}}{\partial x^\mu} V^\mu$$

↑
more general
than Lorentz
transforms
they can be non-linear

Commutator $[X, Y]$ - $X + Y$ are vector fields

linear operator $\Rightarrow [X, Y](f) \equiv X(Y(f)) - Y(X(f))$

vector fields map smooth functions to smooth functions all over the manifold by taking a derivative at every point

$$[X, Y]^\mu = X^\lambda \partial_\lambda Y^\mu - Y^\lambda \partial_\lambda X^\mu$$

Note: partials commute ∂_μ

Tensors (again)

$\omega: T_p \rightarrow \mathbb{R}$ - maps tangent space
to a set of real nos

df - gradient of f , is a one-form

$d/d\lambda$ - vector

$$\therefore df\left(\frac{d}{d\lambda}\right) = \frac{df}{d\lambda} - \text{scalar}$$

The gradients of coords form a natural basis for T_p^* b.c. $\hat{\theta}^{(n)}(\hat{e}_{(r)}) = \delta_r^n$

$$\therefore dx^m(\partial_r) = \frac{\partial x^m}{\partial x^r} = \delta_r^m$$

$$\omega = \omega_m dx^m$$

$$dx^{m'} = \frac{\partial x^{m'}}{\partial x^m} dx^m$$

$$\Rightarrow \omega_{p'} = \frac{\partial x^m}{\partial x^{m'}} \omega_m$$

$$T^{m_1 \dots m_k}_{v_1 \dots v_k} = T(dx^{m_1}, \dots, dx^{m_k}, \partial_{v_1}, \dots, \partial_{v_k})$$

$$\therefore T^{m'_1 \dots m'_k}_{v'_1 \dots v'_k} = \frac{\partial x^{m_1}}{\partial x^{m'_1}} \dots \frac{\partial x^{m_k}}{\partial x^{m'_k}} \frac{\partial x^{v_1}}{\partial x^{v'_1}} \dots \frac{\partial x^{v_k}}{\partial x^{v'_k}} T^{m_1 \dots m_k}_{v_1 \dots v_k}$$

Example

(0,2) Symmetric tensor $S_{\mu\nu} = \begin{pmatrix} 1 & 0 \\ 0 & x^2 \end{pmatrix}$

$$S = S_{\mu\nu} (dx^\mu \otimes dx^\nu)$$

$$= (dx)^2 + x^2(dy)^2$$

choose $x' = \frac{2x}{y}$ $y' = \frac{y}{2}$ where $x > 0 + y > 0$

$$x = \frac{1}{2}x'y \quad y = 2y', \quad x = x'y'$$

$$dx = y'dx' + x'dy'$$

$$dy = 2dy'$$

$$\Rightarrow S = (y'dx' + x'dy')^2 + (x'y')^2(2dy')^2$$

$$\therefore S_{\mu'\nu'} = \begin{pmatrix} y'^2 & x'y' \\ x'y' & x'^2 + 4(x'y')^2 \end{pmatrix}$$

Most tensors are the same under contraction, symmetrization etc. . .

Exceptions are the metric, partial derivatives and Levi-Civita

for partial derivatives

$$\partial_\mu W_\nu = \frac{\partial}{\partial x^{\mu'}} W_{\nu'} = \frac{\partial x^{\mu'}}{\partial x^\mu} \frac{\partial}{\partial x^{\mu'}} \left(\frac{\partial x^{\nu'}}{\partial x^\nu} W_\nu \right)$$
$$= \frac{\partial x^{\mu'}}{\partial x^\mu} \frac{\partial x^{\nu'}}{\partial x^{\nu'}} \left(\frac{\partial}{\partial x^\mu} W_\nu \right)$$

$$+ W_\nu \frac{\partial x^{\mu'}}{\partial x^\mu} \frac{\partial}{\partial x^{\mu'}} \frac{\partial x^{\nu'}}{\partial x^{\nu'}}$$

↑
depends on the partial
deriv of the transformation
Matrix $\therefore$ not a tensor
doesn't transform as a tensor

The Metric

symmetric $(0,2)$ tensor usually nondegenerate

$$g^{\mu\nu} g_{\nu\sigma} = g_{\nu\sigma} g^{\mu\nu} = \delta^\mu_\sigma$$

raises + lowers indices

other uses

- 1) supplies a notion of "past" and "future"
- 2) allows the computation of path length + proper time
- 3) determines the "shortest distance" between pts
- 4) replaces the Newtonian gravitational field
- 5) provides a notion of locally inertial (non rotating) frames
- 6) determines causality
- 7) replaces the dot-product

path length in flat space $ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$

path length in general space $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$

in cartesian coords

$$ds^2 = dx^2 + dy^2 + dz^2$$

in spherical coords

$$x = r \sin\theta \cos\phi$$

$$y = r \sin\theta \sin\phi$$

$$z = r \cos\theta$$

$$ds^2 = dr^2 + r^2 d\theta^2 + r^2 \sin^2\theta d\phi^2$$

$$\Rightarrow g_{\mu\nu} V^\mu W^\nu = g(V, W) = ds^2(V, W)$$

Signature - # of positive + negative eigenvalues
(-, +, +, +)

GR + SR

Euclidean or Riemannian - all positive

→ Lorentzian or pseudo-Riemannian - one minus
indefinite - some +1, some -1

Locally inertial coords - $g_{\hat{\mu}\hat{\nu}}(p) = \eta_{\hat{\mu}\hat{\nu}}$

Local Lorentz frame

$$\partial_{\hat{\sigma}} g_{\hat{\mu}\hat{\nu}}(p) = 0$$

$$\partial_{\hat{\sigma}} \partial_{\hat{\sigma}} g_{\hat{\mu}\hat{\nu}}(p) \neq 0$$

$$g^{\hat{\mu}\hat{\nu}} = \frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \frac{\partial x^\nu}{\partial x^{\hat{\nu}}} g_{\mu\nu}$$

expand in Taylor series

$$\begin{aligned} x^\mu &= \left(\frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \right)_p x^{\hat{\mu}} + \frac{1}{2} \left(\frac{\partial^2 x^\mu}{\partial x^{\hat{\mu}} \partial x^{\hat{\nu}}} \right)_p x^{\hat{\mu}} x^{\hat{\nu}} \\ &+ \frac{1}{6} \left(\frac{\partial^3 x^\mu}{\partial x^{\hat{\mu}} \partial x^{\hat{\nu}} \partial x^{\hat{\omega}}} \right)_p x^{\hat{\mu}} x^{\hat{\nu}} x^{\hat{\omega}} + \dots \end{aligned}$$

assume $x^\mu(p) = x^{\hat{\mu}}(p) = 0$ for simplicity

$$\begin{aligned} \Rightarrow (g^{\hat{\mu}\hat{\nu}})_p &+ (\partial^{\hat{\mu}} g^{\hat{\nu}})_p \hat{x}^{\hat{\mu}} + (\partial^{\hat{\mu}} \partial^{\hat{\nu}} g^{\hat{\rho}})_p \hat{x}^{\hat{\mu}} \hat{x}^{\hat{\nu}} = \left(\frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \frac{\partial x^\nu}{\partial x^{\hat{\nu}}} g \right)_p \\ &+ \left(\frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \frac{\partial^2 x^\nu}{\partial x^{\hat{\mu}} \partial x^{\hat{\nu}}} g + \frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \frac{\partial x^\nu}{\partial x^{\hat{\mu}}} \partial^{\hat{\mu}} g \right)_p \hat{x}^{\hat{\mu}} \\ &+ \left(\frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \frac{\partial^3 x^\nu}{\partial x^{\hat{\mu}} \partial x^{\hat{\nu}} \partial x^{\hat{\omega}}} g + \frac{\partial^2 x^\mu}{\partial x^{\hat{\mu}} \partial x^{\hat{\nu}}} \frac{\partial x^\nu}{\partial x^{\hat{\mu}}} g \right. \\ &\left. + \frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \frac{\partial^2 x^\nu}{\partial x^{\hat{\mu}} \partial x^{\hat{\nu}}} \partial^{\hat{\mu}} g + \frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \frac{\partial x^\nu}{\partial x^{\hat{\mu}}} \partial^{\hat{\mu}} \partial^{\hat{\nu}} g \right)_p \hat{x}^{\hat{\mu}} \hat{x}^{\hat{\nu}} \end{aligned}$$

Set terms of equal $\hat{x}$ equal to each other

$\therefore g^{\hat{\mu}\hat{\nu}}(p)$ are determined by $\left(\frac{\partial x^\mu}{\partial x^{\hat{\mu}}} \right)_p$ — 16 total terms + $\therefore$ possible

set $\partial^{\hat{\mu}} g = 0$ — 40 total terms on RHS

$\therefore$ possible

Example:

What does an observer measure as ordinary 3-velocity?

transform to locally inertial coords

set the observer in the rest frame

$$U^{\hat{\mu}} = (1, 0, 0, 0) - \text{observer}$$

$$V^{\hat{\mu}} = (\gamma, \gamma v, 0, 0) - \text{rocket}$$

$$\gamma = 1/\sqrt{1-v^2} \Rightarrow v = \sqrt{1-\gamma^{-2}}$$

$$\gamma = -g_{\hat{\mu}\hat{\nu}} U^{\hat{\mu}} V^{\hat{\nu}} = -U_{\hat{\mu}} V^{\hat{\mu}}$$

$$\therefore v = \sqrt{1 - (U_{\hat{\mu}} V^{\hat{\mu}})^{-2}}$$

convert back to curved space

$$\Rightarrow v = \sqrt{1 - (U_{\mu} V^{\mu})^{-2}}$$

↖ contraction is the same in either coords